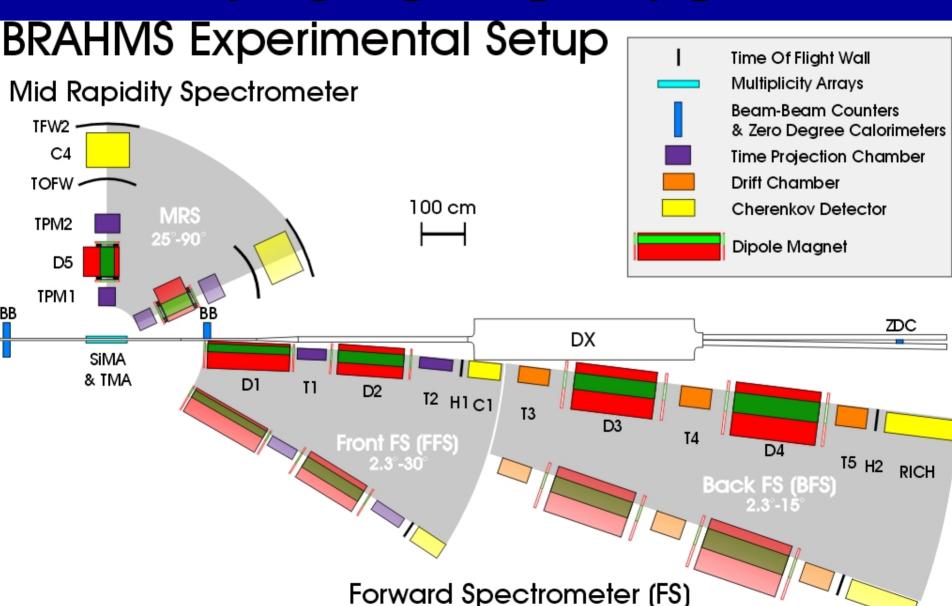
BRAHMS Scanning the phases of QCD

I. Arsene, I.G. Bearden, D. Beavis, C. Besliu, B. Budick, H. Bøggild, C. Chasman, C. H. Christensen, P. Christiansen, J. Cibor, R. Debbe, E. Enger J. J. Gaardhøje, M. Germinario, K. Hagel, O. Hansen, A. Holme, A.K. Holme, H. Ito, A. Jipa, J.I. Jørdre, F. Jundt, C.E.Jørgensen, R. Karabowicz, E.J. Kim, T. Kozik, T.M. Larsen, J.H. Lee, Y. K.Lee, S. Linda, G. Løvhøjden, R. Lystad, Z. Majka, A. Makeev, M. Mikelsen, M. Murray, J. Natowitz, B. Neumann, B.S. Nielsen, K. Olchanski, D. Ouerdane, R.Planeta, F. Rami, C. Ristea, O. Ristea, D.Röhrich, B. H. Samset, D. Sandberg, S. J. Sanders, R.A.Sheetz, P. Staszel, T.S. Tveter, F.Videbæk, R. Wada, Z. Yin and I. S. Zgura

Brookhaven, Strasbourg, Krakov, Johns Hopkins, NYU, Niels Bohr. Texas A&M, Bergen, Bucharest, Kansas, Oslo

Identifying high rapidity particles



Why study rapidity dependence?

- For AuAu we want to understand the limits of jet quenching?
- We also hope to understand the longitudinal dynamics of the source and measure the total energy loss, multiplicity strangeness etc.
- We may find that there is more than one source, i.e. different parts of the system lose causal contact.
- For dAu high rapidity allows us to study the Au nucleus with a faster probe $x = e^{-y} m_T / \sqrt{S}$

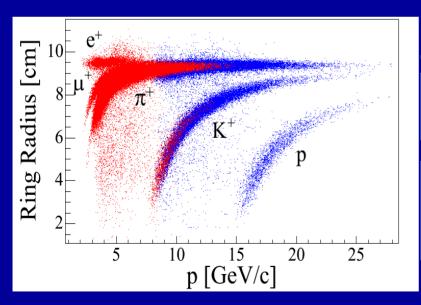
Particle Identification

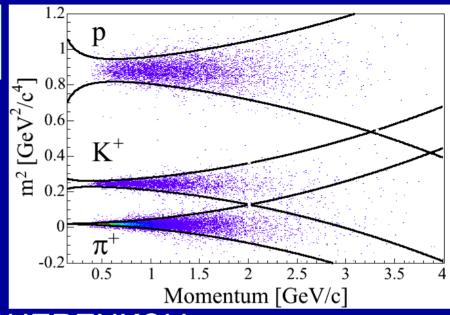
$TIME-OF-FLIGHT m^2 = p^2 \left(\frac{c^2 TOF^2}{r^2}\right)$

$$m^2 = p^2 \left(\frac{c^2 TOF^2}{L^2} - 1 \right)$$

Particle Separation: p_{max} (2 σ cut)=

2σ cut	TOFW	TOF1	TOF2
π / K	2 GeV/c	3 GeV/c	4.5 GeV/c
K/p	3.5GeV/c	5.5GeV/c	7.5GeV/c





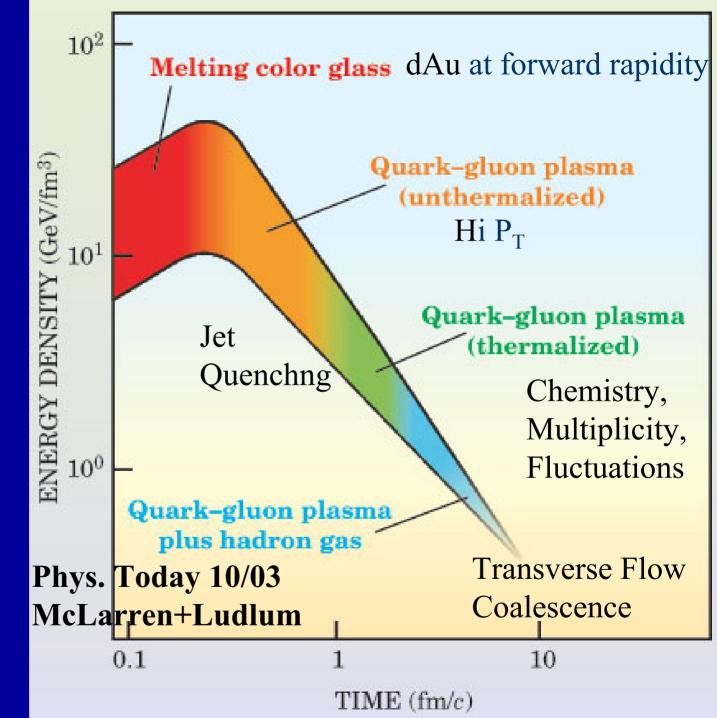
CHERENKO

RICH: Cherenkov light focused on spherical mirror → ring on image plane

Ring radius vs..... momentum gives PID π / K separation 20 GeV/c Proton ID up to 35 GeV/c

A possible time line for Au+Au collisions. This curve itself is a function of rapidity.

This talk will try to run time backwards



What questions can we attack?

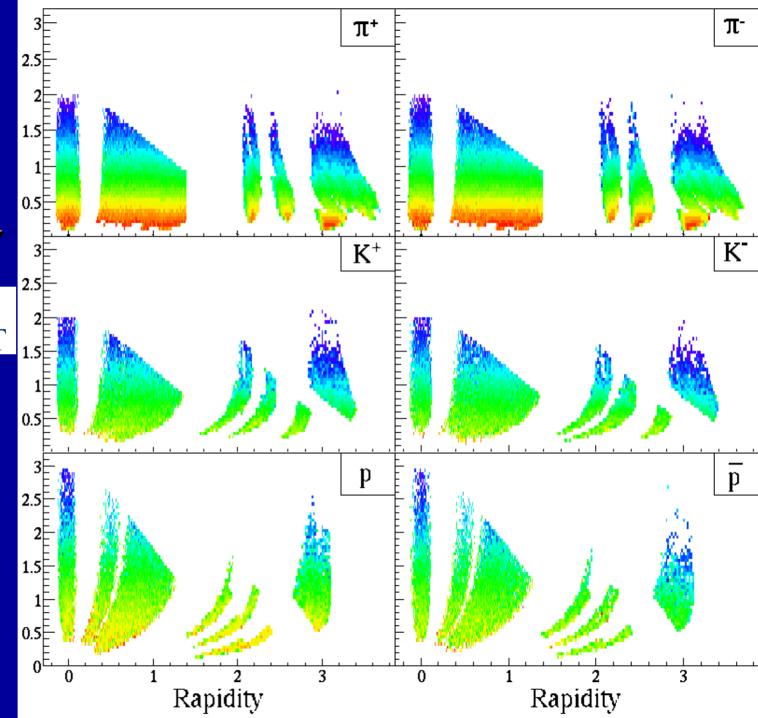
- How much energy is available for particles?
- How many particles?
- What is the volume?
- How do particles flow in the transverse & longitudinal direction?

- What is the chemistry of the system?
- What is the rapidity dependence jet quenching
- What is the nature of the Au wave function at small Feynman x?

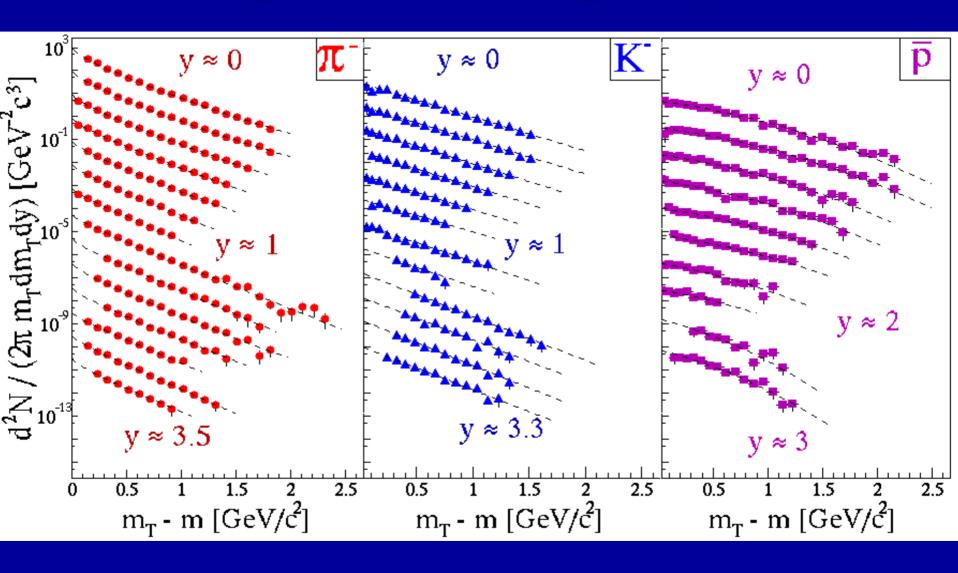
Spectra
vs.
Pt and
rapidity

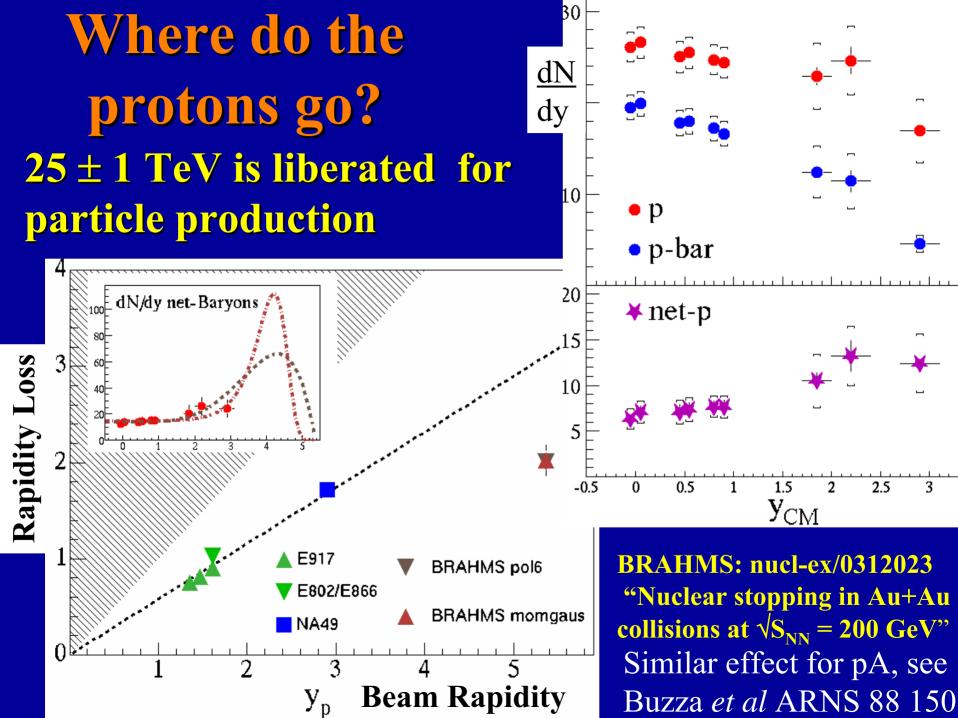
AuAu 10% Central

D. Ouerdane Hadron Spec. Thursday



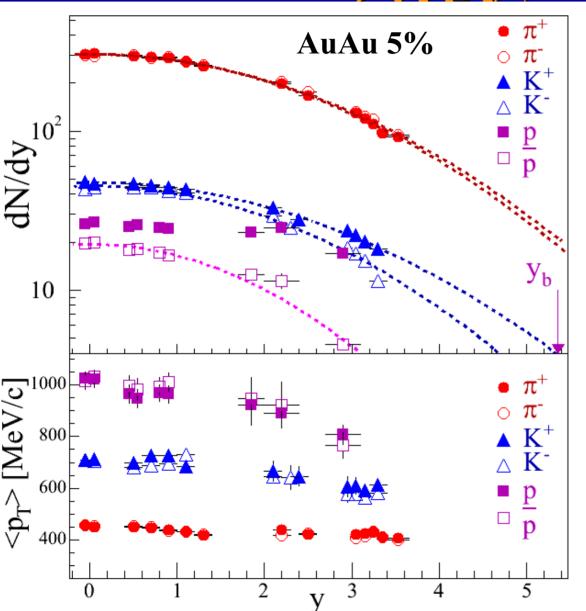
Spectra vs. Rapidity





What can you buy for

25TAV?



	+	_
π	1780	1760
K	290	240
P	85	85

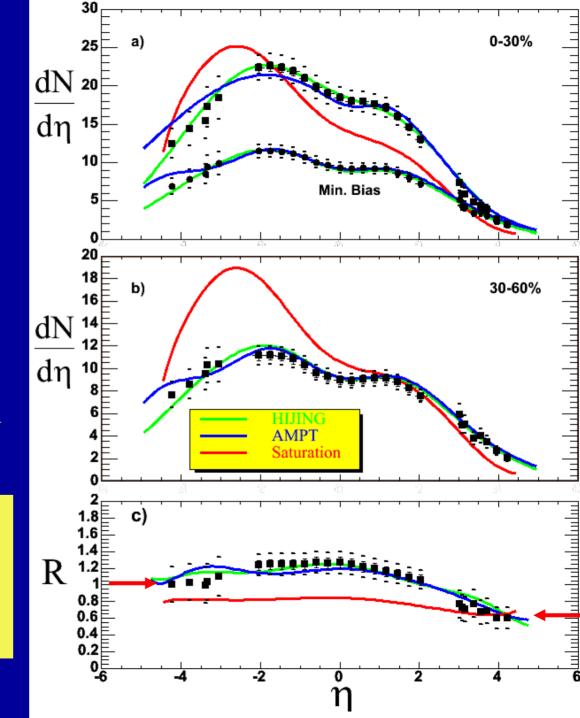
Integrating the energy seen in π , k & P and estimating other particles gives 25 ± 5 TeV

dAu Multiplicity

Data show peaks at η =-1.9 and η =+1.1.

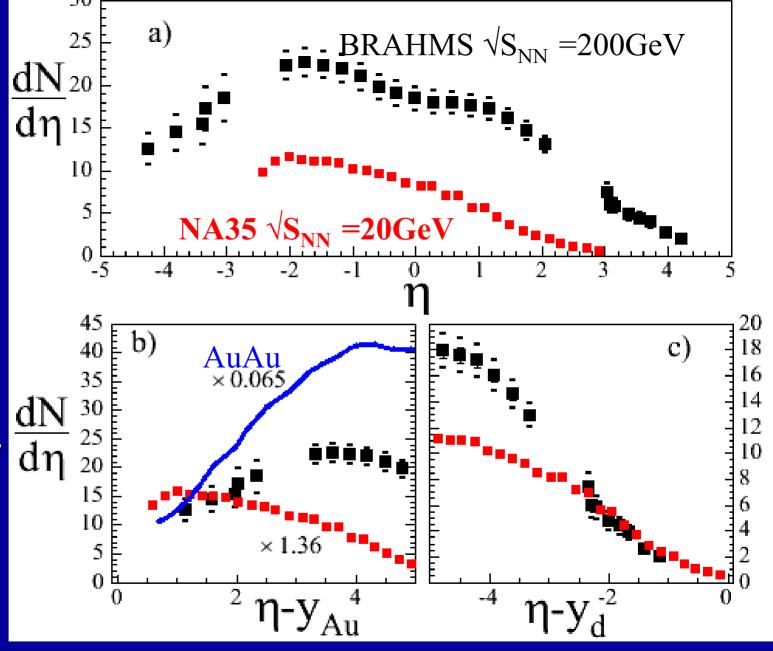
In Au and d regions dN/dη scales with the number of beam participants.

HIJING, & AMPT do well. Saturation model has problems.



dAu
dN/dη
v √S

H. Ito Poster Spectra 18 :

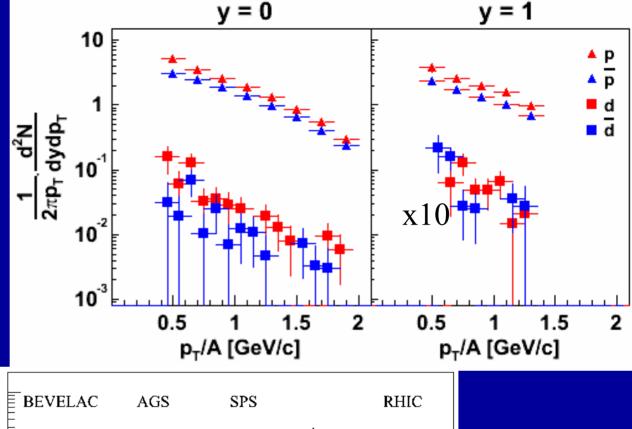


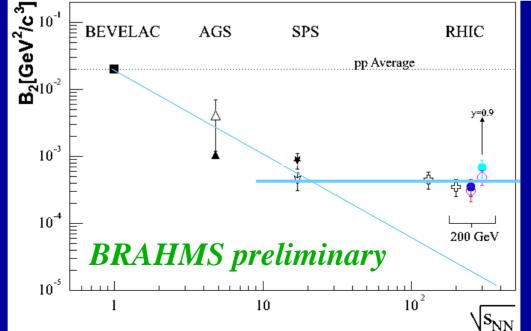
Gold Frame

Deuteron Frame

Coalescence vs. Rapidity

- Compare p, d at same velocity
- Deuterons coalescence from nearby nucleons
- $\mathbf{B}_2 = \mathbf{d}/\mathbf{p}^2$
- $B_2 \propto 1/\text{volume}$





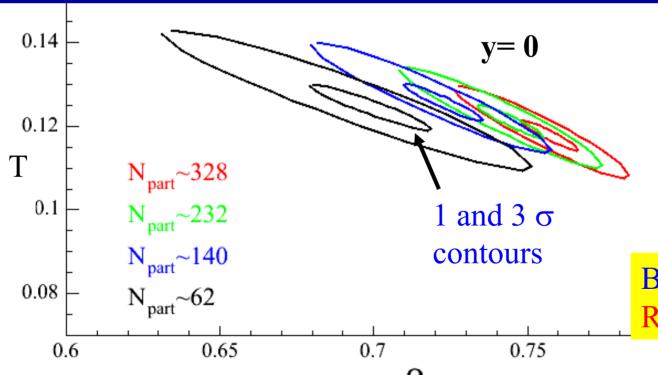
Gemanio + Bearden Poster S27

Flow vs. Centrality

Fit AuAu spectra to blast wave model:

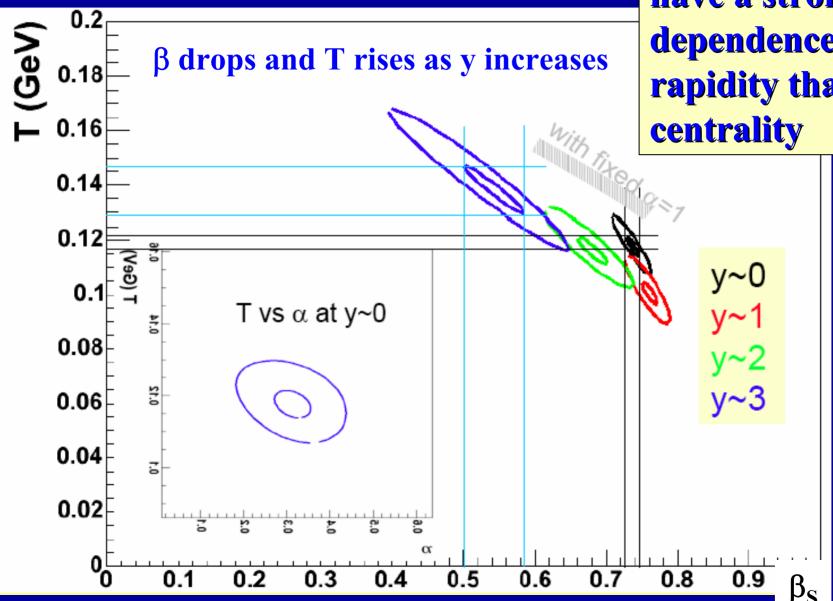
- β_S (surface velocity) drops with dN/d η
- T (temperature) almost constant.

E.J. Kim Poster Spectra 19



Blue y=0 Red y=1p_T (GeV/c)

Flow vs. Rapidity



Flow parameters have a stronger dependence on rapidity than on centrality

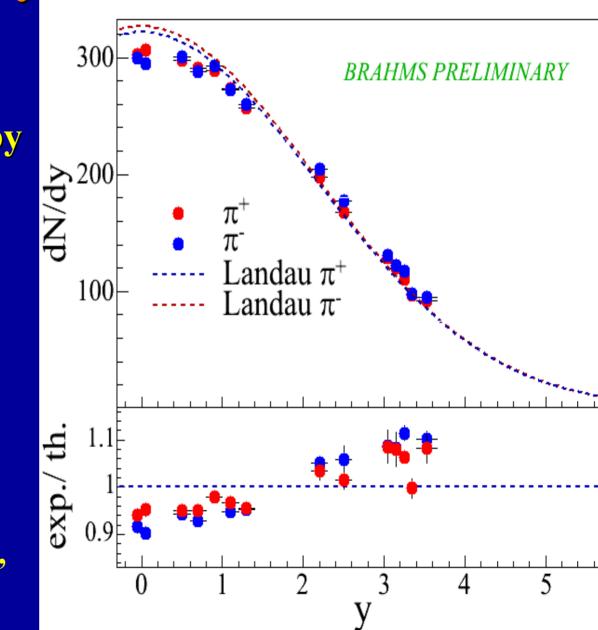
Landau hydrodynamics along beam axis

Assumptions:

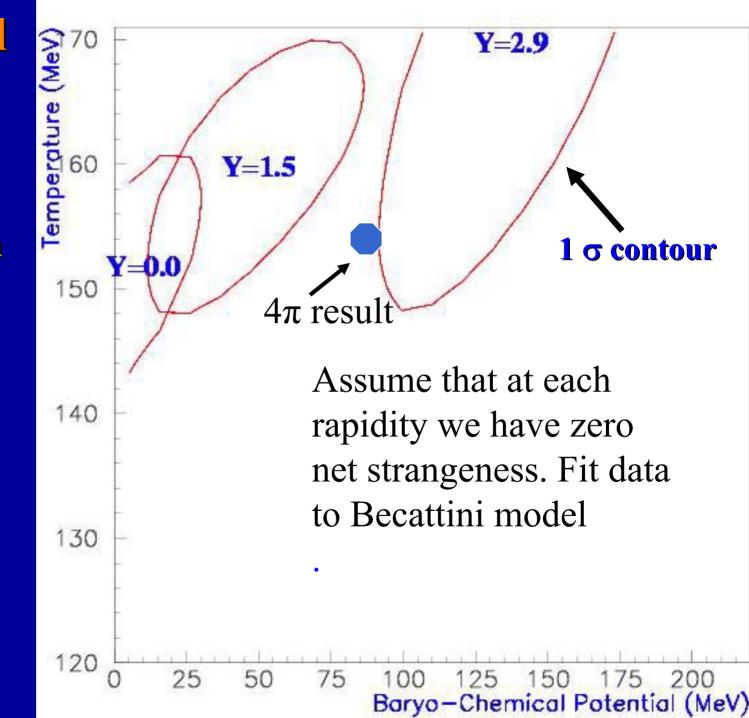
- Isentropic expansion driven by equation of state
- Mass-less particles
- Pt and rapidity factorize

Implications:

- dN/dy Gaussian
- $\sigma = \log (\sqrt{S_{NN}}/2m_p)$ $\approx \log (\gamma_{beam})$
- Model consistent with "limiting fragmentation"



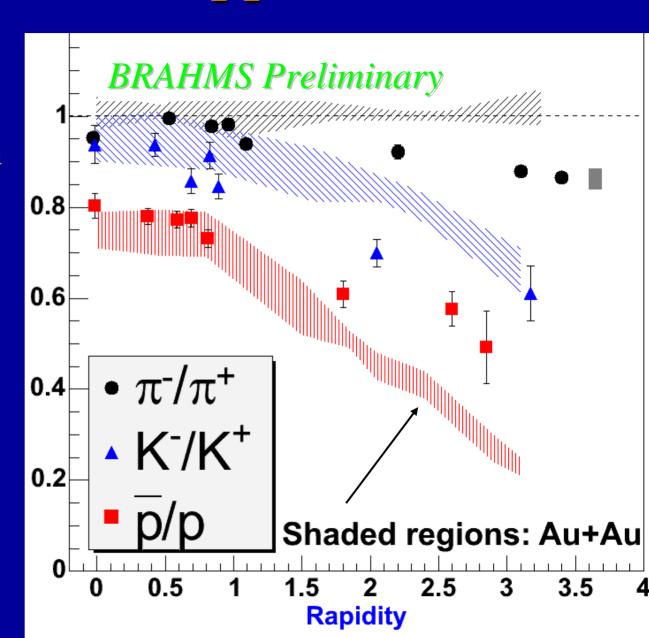
Chemical analysis VS. rapidity
The system may have fewer degrees of freedom at high rapidity J.I Jørdre, **Poster Strange 15**



Particle ratios in pp vs. AuAu

B. H. Samset Poster Spec. 34

pbar/p higher than AuAu while k⁻/k⁺ and π^-/π^+ are lower in pp **Conservation of** isospin and strangeness more important for pp. Stopping different.

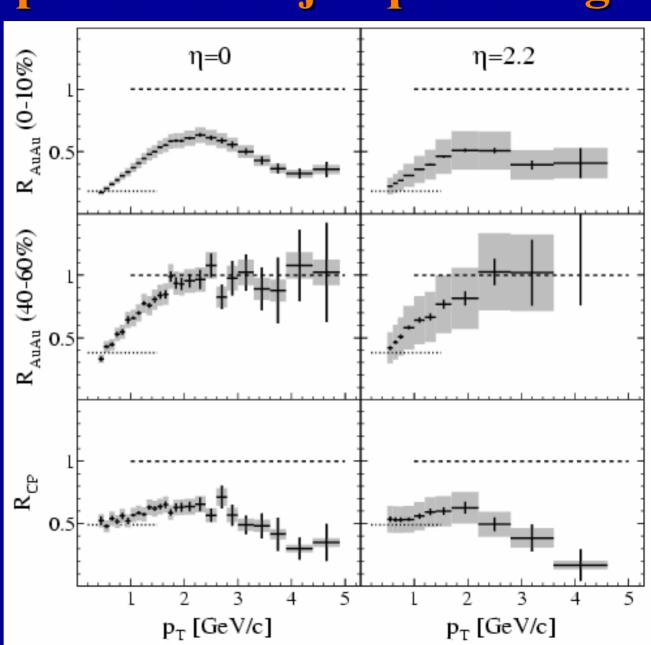


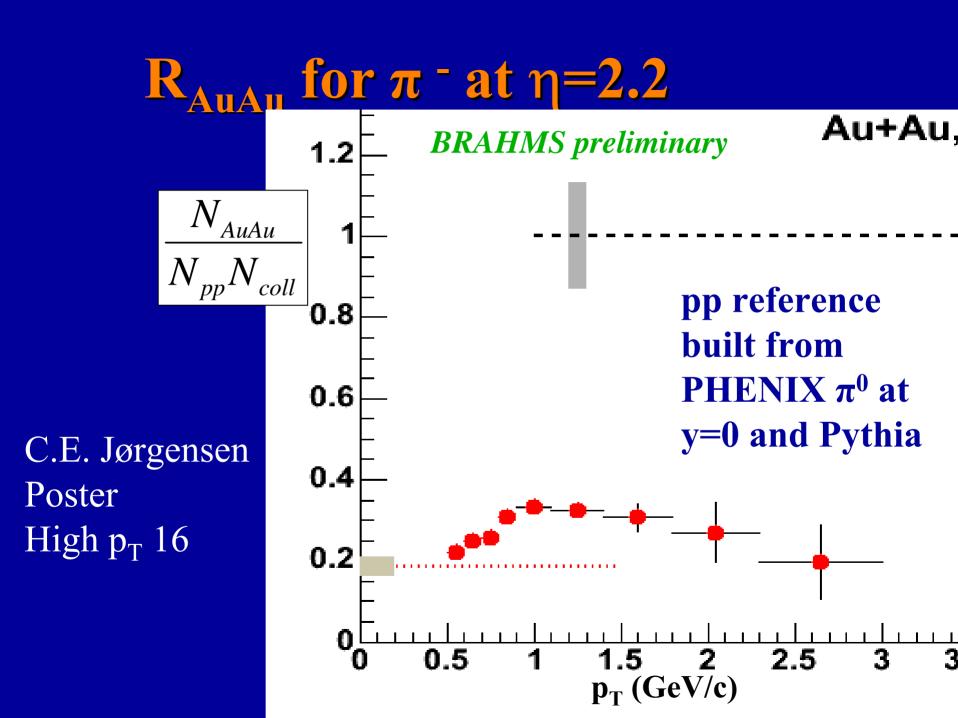
Rapidity dependence of jet quenching

$$R_{AuAu} = \frac{N_{AuAu}}{N_{pp}N_{coll}}$$

- R_{AuAu} < 1 for central collisions
- R_{AuAu} at η=0 and η=2.2 are similar

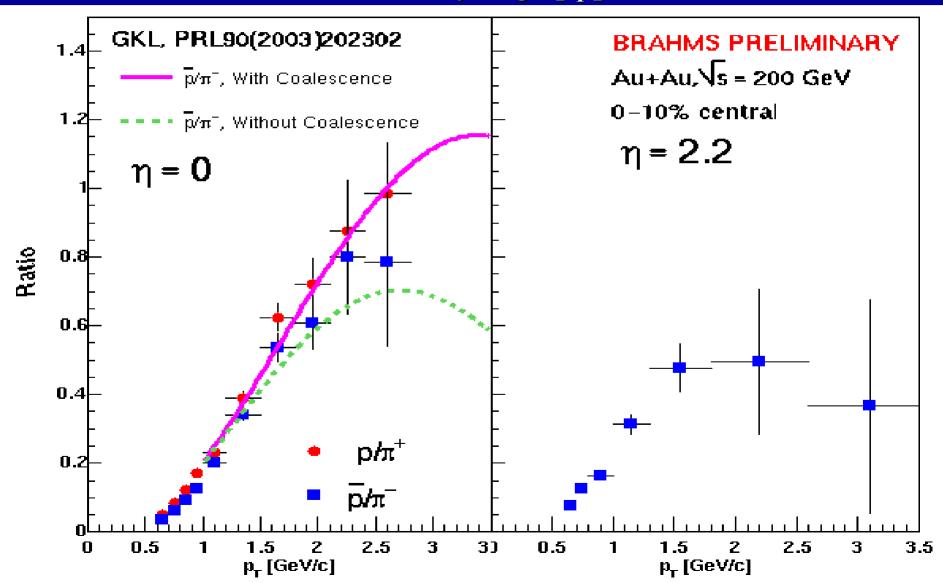
BRAHMS: PRL 91072305 (2003).



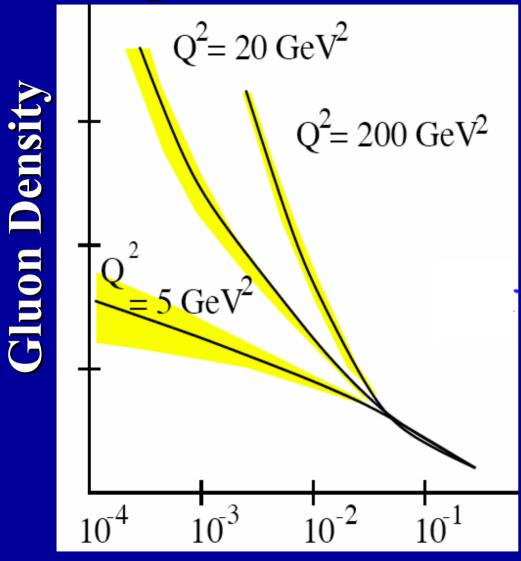


Ratios: anti-protons to pions

Z. Yin Tuesday High p_T parallel



Deep inelastic scattering at HERA

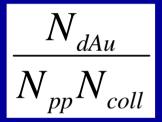


Momentum Fraction x

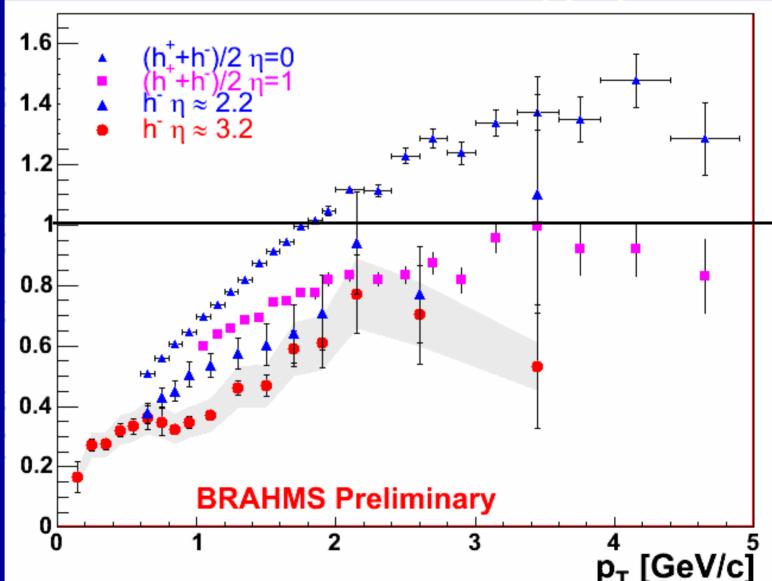
HERA measures the density of gluons as a function of the momentum transfer Q and the gluon's momentum fraction x. For small x there is a universal density function $\mathbf{D} = \mathbf{D} (\mathbf{x}, \mathbf{Q}/\mathbf{Q}_{\mathbf{S}})$ **Q**_S is the saturation scale below which gluons fuse.

Look at a nucleus at high speed

Take the normalized ratio of dAu and pp spectra



R. Debbe Tuesday morning



Moving Forward

- Keep pushing to higher p_T and rapidity with more data and better particle identification.
- Add reaction plane as another "axis" to each measurement.
- Extend coalescence to higher rapidity
- Get more data for dAu (and Au d) ASAP
- Design new experiment for forward region.

What has rapidity taught us?

- $\delta E = 25\pm 1$ TeV in AuAu, broken scaling.
- dAu multiplicity follows beam participant scaling
- Landau hydrodynamics gives rough rapidity dependence of particle production.
- Chemical + kinetic temperatures increase slightly with y, maybe fewer degrees of freedom.
- Jet quenching persists at least to $\eta = 2.2$.
- R_{dAu} at y:0 => 3 suggests gluon saturation.

Backups

Systematics of thermal fits

